



PARTICLE PHYSICS 2011





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RADIATIVE CORRECTIONS

As a rule - size of radiative corrections to a given process is determined by discrepancy between various mass & energy scales involved For a wide class of low-energy and Z-boson observables dominant effects originate entirely in gauge boson propagators (oblique corrections) and can be parametrized in terms of 4 electroweak parameters: $\Delta lpha, \Delta
ho, \Delta r,$ and $\Delta \kappa$ $\mathbf{1}^{\bullet}\Delta \alpha$ determines running fine structure constant at Z boson scale $\alpha(m_Z)/\alpha = (1 - \Delta \alpha)^{-1}$ 2 $\Delta
ho$ measures quantum corrections of NC/CC amplitudes at low energy 3 Δr embodies non-photonic corrections to muon lifetime 4 $\Delta\kappa$ controls effective weak mixing angle $\sin^2 \bar{\theta}_w = \sin^2 \theta_w (1 + \Delta \kappa)$ that occurs in ratio of Zff vector and axial-vector couplings i.e. $c_V^f / c_A^f = 1 - 4|Q_f| \sin^2 \theta_w$

CHESS NOTATION

Today's class — intro to theory of electroweak radiative corrections and its role in testing SM, predicting top mass, constraining Higgs mass, and searching for deviations that may signal presence of new physics Implementing such a program can be first formulated from point of view of experimentalist

Introducing notation – $\sin^2 \theta_w = s^2 = 1 - c^2, \ m_W^2 \equiv w, \ m_Z^2 \equiv z$

Electroweak theory predicts at Born level that:

$$\frac{w}{z} = 1 - s^2$$

$$\frac{\pi\alpha}{\sqrt{2}G_F}\frac{1}{w} = s^2$$

$$\frac{\sigma(\nu_{\mu}e)}{\sigma(\bar{\nu}_{\mu}e)} = \frac{3 - 12s^2 + 16s^4}{1 - 4s^2 + 16s^4}$$

$$\frac{\Gamma(Z \to f\bar{f})}{m_Z} = \frac{\alpha}{3} C_F \left((c_V^f)^2 + (c_A^f)^2 \right)$$

$$A_{\rm LR} \simeq A_{\tau} \simeq \left[\frac{4}{3}A_{\rm FB}\right]^{1/2} \simeq 2(1-4s^2)$$



Chess Eqs. represent incomplete list of experiments capable of measuring $\sin^2 heta_w$ Validity of SM requires that each measurement yields same value of s^2 Ratio of u_{μ} scattering on left- and right-handed electrons which is a function of $\sin^2 \theta_w$ only -1| Measurement of weak boson masses - 🚊 II Combination of $m_W, lpha$, and G_F as determined by muon lifetime $rac{1}{2}$ IV Partial widths of Z into a fermion pair with vector and axial coupling c_V^{f} and c_A^{f} and color factor $C_F=3~(1)$ for quarks (leptons) -V Various asymmetries measured at Z-factories 🖛 👑

CORRECTIONS TO Z PROPAGATOR

After inclusion of $\mathcal{O}(\alpha)$ corrections $\sin^2 \theta_w$ values obtained from different methods will no longer be same because radiative corrections modify chess eqs. in different ways



It does not -- however -- contribute to $\mathcal{O}(lpha)$ shifts in W,Z masses



which yield an improved $\sin^2 \theta_w$ value via \blacksquare

WEINBERG ANGLE

Experimentalist has to make a choice and define Weinberg angle to $\mathcal{O}(\alpha)$ All other experiments should be reformulated in terms of preferred s^2 What this choice should be is no longer a matter of debate we will define $\sin^2 \theta_w$ in terms of physical masses of weak bosons

e.
$$\sin^2 \theta_w \equiv 1 - \frac{m_W^2}{m_Z^2} = 0.23122(15)$$
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This choice is particularly useful in that one can estimate radiative corrections in terms of renormalization group which has been previously introduced

 $\mathcal{O}(\alpha)$ corrections can be qualitatively understood in terms of loop corrections to vector-boson propagators \mathcal{T} and \clubsuit

A most straightforward test of theory is now obtained by fixing $\sin^2 \theta_w$ in terms of measured weak boson masses and verifying that its value coincides with other measurements of θ_w

FREE PARAMETERS

All UV divergences in QED can be absorbed in two parameters: α and m_e List of parameters to be fixed by experiment in electroweak theory includes

 $\alpha, m_W, m_Z, m_H, m_f$

weak mixing angle does not appear in list of parameters: its value is automatically determined by m_W , m_Z via \Im

Traditionally - SM Lagrangian is determined in terms of

$$g, g', \lambda, \mu, Y_f$$

There is in fact a direct translation between sets 🖑 and 🖑

$$g^2 = e^2 \frac{z}{z - w}$$

$$\lambda = e^2 \frac{z m_H^2}{8w(z-w)}$$

$$g'^2 = e^2 \frac{z}{w}$$

$$Y_f = e^2 \frac{zm_f^2}{2w(z-w)}$$

ELECTROMAGNETIC RADIATIVE CORRECTIONS

EXAMPLE - Show how relation 1 is calculated to $\mathcal{O}(\alpha)$ in terms of weak angle θ_w defined by 1

Origin of relation 1 is μ lifetime - to leading order is given by diagram



In Fermi theory \blacktriangleright electromagnetic radiative corrections must be included to obtain result to $\mathcal{O}(\alpha)$



ELECTROWEAK RADIATIVE CORRECTIONS

In electroweak theory



ELECTROWEAK MAODEL AT BORN LEVEL

Equating
$$\checkmark$$
 and \digamma we obtain $G_F = \frac{\pi \alpha}{\sqrt{2}} \frac{1}{ws^2} (1 + \Delta r) \implies$
with $\Delta r = \Delta \alpha - \frac{c^2}{s^2} \Delta \rho + \Delta_{\text{rem}} \bigcirc$
efore discussing status of measurements of $\Delta r =$ note that
to leading order $\Rightarrow \Delta r = 0$

$$\alpha = \frac{e^2}{4\pi} \qquad m_W = \frac{g v}{2} = \frac{g}{2\sqrt{2\lambda}} m_H \qquad \text{and} \qquad m_Z = \frac{m_W}{\cos \theta_w}$$



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DOMNINANT TERMS

Electroweak radiative corrections are gathered in Δr

We specifically isolated fermions which are responsible for running of



as well as third generation heavy quark diagram



Other contributions are small and are grouped in remainder $\Delta_{
m rem}$. To the extent that $\Delta_{
m rem}$ is small — one can imagine summing series

$$m \longrightarrow + m \longrightarrow + \cdots$$

$$eplacement \ (1 + \Delta r) \rightarrow (1 - \Delta r)^{-1} \quad in \$$

We already discussed running of lpha from small lepton masses to m_Z $\Delta \alpha = \sum_{f} \mathbf{\gamma} \mathbf{\gamma} \mathbf{\gamma}$ $\Delta \rho = \underbrace{W}_{W}$ Other large contribution $\Delta
ho$ which represents loop is our primary focus here Its value is given by $\Delta \rho = \frac{\alpha}{4\pi} \frac{z}{\omega(z-\omega)} N_C |U_{tb}|^2 \left[m_t^2 F(m_t^2, m_b^2) + m_b^2 F(m_b^2, m_t^2) \right] \clubsuit$ with $F(m_1^2, m_2^2) = \int_0^1 dx \, x \, \ln\left[m_1^2(1-x) + m_2^2 x\right]$ where $N_C=3$ is number of colors and U_{tb} is CKM matrix element $|U_{tb}|^2 \simeq 1$

FINGEPRINTS OF ELECTROWEAK THEORY

Diagram
$$\Delta \rho = \sum_{W} \bigcup_{\overline{b}}^{t} W$$

has important property that defining
$$m_t=m_b+\epsilon$$
 .



In QED

only equal mass fermions and antifermions appear in neutral photon loops and diagrams of this type are not possible

Rewrite and * in form

$$\begin{split} \Delta \rho &= \frac{G_F}{4\pi} \left[m_t^2 + m_b^2 - \frac{2m_b^2 m_t^2}{m_t^2 - m_b^2} \ln \frac{m_t^2}{m_b^2} \right] \\ &\simeq \frac{G_F}{4\pi} m_t^2 \simeq \frac{3\alpha}{16\pi} \frac{1}{c^2 s^2} \frac{m_t^2}{z} \end{split}$$

 m_t^2/z contribution to an observable is forbidden in QED and QCD where virtual particle effects are suppressed by inverse powers of their masses embodies this requirement because $\epsilon = 0$ for photon loops Conversely - appearance of m_t^2/z term is a characteristic feature of electroweak theory WE ARE NOW READY TO ILLUSTRATE THAT $\Delta \rho \neq 0$... We first determine experimental value of Δr from pproxUsing $\frac{G_F}{\sqrt{2}} = \frac{g^2}{8m_W^2}$ and $e = g\sin\theta_w$ $\Delta r_{\rm exp} \simeq 1 - (37.281 \text{ GeV})^2 \frac{z}{\omega(z-\omega)} \simeq 0.035$ We next recall $\Delta \alpha \simeq 1 - \frac{\alpha(0)}{\alpha(m_Z^2)} \simeq 1 - \frac{128}{137} \simeq 0.066$ **I** O Crucial point is that $\Delta r_{
m exp}
eq \Delta \alpha ~(\ A
eq {f I})$ ${\cal O}(lpha)$ SM relation ${\cal O}$ requires a non-vanishing value of $\Delta
ho$ Using st we obtain that $\Delta
ho=0.0086$ and Θ yields $(\Delta r)_{\text{calculated}} = \Delta \alpha - \frac{c^2}{c^2} \Delta \rho = 0.037$ in agreement with experimental value We leave it as an exercise to insert errors into calculation and show that argument survives a straightforward statistical analysis

CONSTRAINS ON HIGGS MAASS

Higgs particle makes a contribution to Δr



From $114.4 \ {
m GeV} < m_H \lesssim 1 \ {
m TeV}$ we obtain that $\Delta h < 0.0006$ contribution too small to be sensed by simple analysis presented above Quantity Δr is in principle sensitive to Higgs mass

Radiative corrections predicted by SM have successfully confronted experiment program is however far from complete problem can be quantified by rewriting Δ and \bigcirc as

$$\Delta r_{\rm exp} = F(m_W, m_t, m_H)$$

Using Z-pole measurements of SLD and LEP1 electroweak radiative corrections are evaluated to predict masses of top quark and W-boson

(m_W, m_t) plane



Contour curves of 68% CL in (m_t, m_W) plane for direct measurements and indirect determinations Band shows correlation between m_W and m_t expected in SM

NEUTRINO OSCILLATION

Convincing experimental evidence exists

for oscillatory transitions $\nu_{\alpha} \rightleftharpoons \nu_{\beta}$ between different neutrino flavors Simplest and most direct interpretation of atmospheric data is that of muon neutrino oscillations

Evidence of atmospheric u_{μ} disappearing is now at $>15\sigma$ most likely converting to $u_{ au}$

Angular distribution of contained events shows that for $E_{\nu} \sim 1~{
m GeV}$ deficit comes mainly from $L_{
m atm} \sim 10^2-10^4~{
m km}$

These results have been confirmed by KEK-to-Kamioka (K2K) experiment which observes disappearance of accelerator ν_{μ} at a distance of 250 km Data collected by the Sudbury Neutrino Observatory (SNO) in conjunction with data from Super-Kamiokande (SK) show that solar ν_e convert to ν_{μ} or ν_{τ} with CL of more than 7σ

KamLAND Collaboration has measured flux of $\overline{\nu}_e$ from distant reactors and find that $\overline{\nu}_e$ disappear over distances of about 180 kmAll these data suggest that neutrino eigenstates that travel through space are not flavor states that we measured through weak force but rather mass eigenstates

LEPTON FLAVOR MAIXING

Flavor eigenstates $|
u_{lpha}
angle$ and mass eigenstates $|
u_i
angle$ (i.e. mixing matrix) are related by a unitary transformation U (i.e. mixing matrix)

$$\begin{split} |\nu_{\alpha}\rangle &= \sum_{i} U_{\alpha i} |\nu_{i}\rangle \Leftrightarrow |\nu_{i}\rangle = \sum_{\alpha} (U^{\dagger})_{i\alpha} |\nu_{\alpha}\rangle = \sum_{\alpha} U_{\alpha i}^{*} |\nu_{\alpha}\rangle \\ \text{sith} \quad U^{\dagger}U = \mathbb{I}, \text{ i.e., } \sum_{i} U_{\alpha i} U_{\beta i}^{*} = \delta_{\alpha\beta} \quad \text{and} \quad \sum_{\alpha} U_{\alpha i} U_{\alpha j}^{*} = \delta_{ij} \end{split}$$

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For antineutrinos one has to replace $U_{lpha i}$ by $U_{lpha i}^{*}$

.e.
$$\left| \bar{\nu}_{\alpha} \right\rangle = \sum_{i} U^{*}_{\alpha i} \left| \bar{\nu}_{i} \right\rangle$$

Number of parameters of an $n \times n$ unitary matrix is 2nIt is easy to see that 2n - 1 relative phases of n^2 neutrino states can be redefined such that $(n - 1)^2$ independent parameters are left For these it is convenient to take:

 $\frac{1}{2}n(n-1)$ weak mixing angles of an n-dimensional rotation and $\frac{1}{2}(n-1)(n-2)$ CP-violating phases

HARAILTONIAN MAECHANICS

ullet Being eigenstates of mass matrix ullet states $|
u_i
angle$ are stationary states

(i.e. they have the time dependence)

$$|\nu_i(t)\rangle = e^{-iE_it}|\nu_i\rangle$$

with
$$E_i = \sqrt{p^2 + m_i^2} \approx p + \frac{m_i^2}{2p} \approx E + \frac{m_i^2}{2E}$$

where Epprox p is total neutrino energy here it is assumed that neutrinos are stable A pure flavor state $|
u_{lpha}
angle = \sum U_{lpha i} |
u_i
angle$ present at t=0

develops with time into state

$$|\nu(t)\rangle = \sum_{i} U_{\alpha i} e^{-iE_{i}t} |\nu_{i}\rangle = \sum_{i,\beta} U_{\alpha i} U_{\beta i}^{*} e^{-iE_{i}t} |\nu_{\beta}\rangle$$

TRANSITION ARAPLITUDE

Time dependent transition amplitude for transition from flavor u_{α} to flavor u_{β} is

$$\begin{aligned} \mathfrak{A}(\nu_{\alpha} \to \nu_{\beta}) &\equiv \langle \nu_{\beta} | \nu(t) \rangle &= \sum_{i} U_{\alpha i} U_{\beta i}^{*} e^{-iE_{i}t} \\ &= \sum_{i,j} U_{\alpha i} \delta_{ij} e^{-iE_{i}t} (U^{\dagger})_{j\beta} \\ &= (UDU^{\dagger})_{\alpha\beta} \end{aligned}$$

 \bigotimes

with $D_{ij} = \delta_{ij} e^{-iE_i t}$ (diagonal matrix) An equivalent expression for transition amplitude is obtained by inserting \Re into \boxtimes and extracting an overall phase factor e^{-iEt}

$$\begin{aligned} \mathfrak{A}(\nu_{\alpha} \to \nu_{\beta}, t) &= \sum_{i} U_{\alpha i} U_{\beta i}^{*} e^{-\frac{i m_{i}^{2} t}{2E}} \\ &= \sum_{i} U_{\alpha i} U_{\beta i}^{*} e^{-\frac{i m_{i}^{2} L}{2E}} \end{aligned}$$

where L=ct (recall c=1) is distance of detector in which u_{eta} is observed from u_{lpha} source

$i \rightarrow j$ TRANSITION ARAPLITUDE

For an arbitrary chosen fixed j transition amplitude becomes

$$\widetilde{\mathfrak{A}}(\nu_{\alpha} \to \nu_{\beta}, t) = e^{iE_{j}t}\mathfrak{A}(\nu_{\alpha} \to \nu_{\beta}, t)$$

$$= \sum_{i} U_{\alpha i} U_{\beta i}^{*} e^{-i(E_{i} - E_{j})} t$$

$$= \delta_{\alpha\beta} + \sum_{i} U_{\alpha i} U_{\beta i}^{*} \left[e^{-i(E_{i} - E_{j})t} - 1 \right]$$

$$= \delta_{\alpha\beta} + \sum_{i \neq j} U_{\alpha i} U_{\beta i}^{*} \left[e^{-i\Delta_{ij}} - 1 \right]$$

with
$$\Delta_{ij} = (E_i - E_j) = 1.27 \frac{\delta m_{ij}^2 L}{E}$$
 when
 L is measured in km E in GeV and $\delta m_{ij}^2 = m_i^2 - m_j^2$ in eV^2
(in T unitarity relation \oplus has been used)

PROPERTIES OF TRANSITION ANAPLITUDE \checkmark Transition amplitudes are thus given by n(n-1) real parameters $(n-1)^2$ independent parameters of unitary matrix (which determines sizes of oscillations) and n-1 mass square differences (which determine frequencies of oscillations) \checkmark If CP is conserved in neutrino oscillations all CP-violating phases vanish and $U_{lpha i}$ are real i.e. U is an orthogonal matrix $U^{-1} = U^T$ with $rac{1}{2}n(n-1)$ parameters Number of parameters for transition amplitude is then $rac{1}{2}(n-1)(n+2)$ 🗸 Using 😌 we obtain amplitudes for transitions between antineutrinos $\mathfrak{A}(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}; t) = \sum_{i} U_{\alpha i}^{*} U_{\beta i} e^{-iE_{i}t}$

CPT

Comparing 🖾 and 🔆 following relation holds for transformations between neutrinos and antineutrinos

$$\mathfrak{A}(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) = \mathfrak{A}(\nu_{\beta} \to \nu_{\alpha}) \neq \mathfrak{A}(\nu_{\alpha} \to \nu_{\beta})$$

This follows directly from CPT theorem: C changes particle into antiparticle P provides necessary flip from left-handed neutrino to right-handed antineutrino and vice versa and T reverses arrow indicating transition If CP is conserved $= U_{\alpha i}$ and $U_{\beta i}$ are real in \boxtimes and \neq That is = if time reversal invariance holds one has

$$\mathfrak{A}(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) = \mathfrak{A}(\nu_{\alpha} \to \nu_{\beta}) = \mathfrak{A}(\bar{\nu}_{\beta} \to \bar{\nu}_{\alpha}) = \mathfrak{A}(\nu_{\beta} \to \nu_{\alpha})$$

Therefore – CP violation can be searched for e.g. by comparing oscillations $\nu_{\alpha} \rightarrow \nu_{\beta}$ and $\nu_{\beta} \rightarrow \nu_{\alpha}$

TRANSITION PROBABILITY

Transition probabilities are obtained by squaring moduli of amplitudes 🛛

$$P_{\nu_{\alpha} \to \nu_{\beta}} = \left| \sum_{i} U_{\alpha i} U_{\beta i}^{*} e^{-iE_{i}t} \right|^{2}$$

$$= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re \left(U_{\alpha i}^{*} U_{\beta i} U_{\alpha j} U_{\beta j}^{*} \right) \sin^{2} \Delta_{ij}$$

$$+ 2 \sum_{i>j} \Im \left(U_{\alpha i}^{*} U_{\beta i} U_{\alpha j} U_{\beta j}^{*} \right) \sin 2\Delta_{ij}$$

MATCHING EXPERIMENTAL DATA In standard treatment of neutrino oscillations flavor eigenstates $|
u_{lpha}
angle(lpha=e,\,\mu,\, au)$ are expanded in terms of 3 mass eigenstates $|
u_i
angle(i=1,\,2,\,3)$ * Atmospheric neutrino data suggest corresponding corresponding oscillation phase must be maximal $\Delta_{
m atm} \sim 1$ which requires $\delta m^2_{
m atm} \sim 10^{-4} - 10^{-2} \ {
m eV}^2$ Assuming all upgoing multi-GeV u_{μ} oscillate into different flavor while none of downgoing ones do observed up-down asymmetry leads to a mixing angle very close to maximal $\sin^2 2 heta_{
m atm} > 0.85$ * Combined analysis of atmospheric neutrinos with K2K leads to a best fit-point and 1σ ranges $\delta m_{\rm atm}^2 = 2.2^{+0.6}_{-0.4} \times 10^{-3} \ {\rm eV}^2$ and $\tan^2 \theta_{\rm atm} = 1^{+0.35}_{-0.26}$ st On the other hand st reactor data suggest $|U_{e3}|^2 \ll 1$

MAASS EIGENSTATES

TO SIMPLIFY ...

we use fact that $\left|U_{e3}\right|^2$ is nearly zero to ignore possible CP violation and assume that elements of U are real

With this in mind - we can define a mass basis as follows

 $|\nu_1\rangle = \sin\theta_{\odot}|\nu^{\star}\rangle + \cos\theta_{\odot}|\nu_e\rangle$

$$|\nu_2\rangle = \cos\theta_{\odot}|\nu^{\star}\rangle - \sin\theta_{\odot}|\nu_e\rangle$$

where $heta_{\odot}$ is solar mixing angle

$$|\nu_3\rangle = \frac{1}{\sqrt{2}}(|\nu_\mu\rangle + |\nu_\tau\rangle)$$

$$\left| |
u^{\star}
ight
angle = rac{1}{\sqrt{2}} (|
u_{\mu}
angle - |
u_{ au}
angle)
ight|$$
 \triangleq is eigenstate orthogonal to $|
u_{3}
angle$

FLAVOR EIGENSTATES

Inversion of neutrino mass-to-flavor mixing matrix leads to

$$|\nu_e\rangle = \cos\theta_{\odot}|\nu_1\rangle - \sin\theta_{\odot}|\nu_2\rangle | \neq |\nu^*\rangle = \sin\theta_{\odot}|\nu_1\rangle + \cos\theta_{\odot}|\nu_2\rangle$$

by adding B and \triangleq one obtains u_{μ} flavor eigenstate

$$|\nu_{\mu}\rangle = \frac{1}{\sqrt{2}} \left[|\nu_{3}\rangle + \sin\theta_{\odot} |\nu_{1}\rangle + \cos\theta_{\odot} |\nu_{2}\rangle \right]$$

and by substracting these same equations ν_{τ} eigenstate Combined analysis of Solar neutrino data and KamLAND data are consistent at 3σ CL with best-fit point and 1σ ranges:

 $\delta m_{\odot}^2 = 8.2^{+0.3}_{-0.3} \times 10^{-5} \text{ eV}^2$

 $\tan^2 \theta_{\odot} = 0.39^{+0.05}_{-0.04}$

FAR-OUT NEUTRINOS

For $\Delta_{ij} \gg 1$ phases will be erased by uncertainties in L and E(as would be case for neutrinos propagating over cosmic distances), averaging over $\sin^2 \Delta_{ij}$ in \mathbf{H} we obtain

$$P(\nu_{\alpha} \to \nu_{\beta}) = \delta_{\alpha\beta} - 2\sum_{i>j} U_{\alpha i} U_{\beta i} U_{\alpha j} U_{\beta j}$$



$$P(\nu_{\alpha} \to \nu_{\beta}) = \delta_{\alpha\beta} - \sum_{i,j} U_{\alpha i} U_{\beta i} U_{\alpha j} U_{\beta j} + \sum_{i} U_{\alpha i} U_{\beta i} U_{\alpha i} U_{\beta i}$$
$$= \delta_{\alpha\beta} - \left(\sum_{i} U_{\alpha i} U_{\beta i}\right)^{2} + \sum_{i} U_{\alpha i}^{2} U_{\beta i}^{2}$$

Since $\delta_{lphaeta} = \delta_{lphaeta}^2$ first and second terms in * cancel each other

 $P(\nu_{\alpha} \to \nu_{\beta}) = \sum_{i} U_{\alpha i}^{2} U_{\beta i}^{2}$

AVERAGE PROBABILITY

Probabilities for flavor oscillation are then

$$P(\nu_{\mu} \to \nu_{\mu}) = P(\nu_{\mu} \to \nu_{\tau}) = \frac{1}{4} \left(\cos^4 \theta_{\odot} + \sin^4 \theta_{\odot} + 1\right)$$

$$P(\nu_{\mu} \to \nu_{e}) = P(\nu_{e} \to \nu_{\mu}) = P(\nu_{e} \to \nu_{\tau}) = \sin^{2} \theta_{\odot} \ \cos^{2} \theta_{\odot}$$

$$P(\nu_e \to \nu_e) = \cos^4 \theta_{\odot} + \sin^4 \theta_{\odot}$$

Let ratios of neutrino flavors at production in cosmic sources be written as $w_e: w_\mu: w_\tau$ with $\sum_{\alpha} w_\alpha = 1$ so that relative fluxes of each mass eigenstate are $w_j = \sum \omega_\alpha U_{\alpha j}^2$

We conclude that probability of measuring on Earth a flavor α is

$$P_{\nu_{\alpha} \text{ detected}} = \sum_{j} U_{\alpha j}^{2} \sum_{\beta} w_{\beta} U_{\beta j}^{2}$$

EARHTLY RATIOS

- Q Any initial flavor ratio that contains $w_e=1/3$ will arrive at Earth with equipartition on 3 flavors
- O Cosmic neutrinos arise dominantly from decay of charged pions their initial flavor ratios of nearly 1:2:0 should arrive at Earth democratically distributed
- $^{oldsymbol{O}}$ Fairly robust prediction of 1:1:1 cosmic neutrino flavor ratios
- O Prediction for pure $\bar{\nu}_e$ source originating via neutron β -decay has different implications for flavor ratios $w_e=1$ yields Earthly ratios $\sim 5:2:2$
- Such a unique ratio would appear above 1:1:1 background in direction of neutron source

O Flavor identification of neutrinos on a statistical basis becomes possible at IceCube opening up a window for discoveries in particle physics not otherwise accessible to experiment

the Yampire



In SM masses arise from Yukawa interactions which couple right-handed fermion with its left-handed doublet after spontaneous symmetry breaking
 Because no right-handed neutrinos exist in SM Yukawa interactions leave neutrinos massless
 One may wonder if neutrino masses could arise from loop corrections or even by non-perturbative effects but this cannot happen because any neutrino mass term that can be constructed with SM fields would violate total lepton symmetry

To introduce a neutrino mass term we must either extend particle content or else abandon gauge invariance and/or renormalizability

How to kill a Yampire

★ We keep gauge symmetry and introduce arbitrary number *m* of additional right-handed neutrino states (singlets under hypercharge) *V*_{Rs}(1,1)₀
★ With particle contents of SM and addition of an arbitrary *m* number of right-handed neutrinos one can construct two types of mass terms that arise from gauge invariant renormalizable operators

$$-\mathcal{L}_{M_{\nu}} = \sum_{\alpha=e,\mu,\tau} \sum_{i=1}^{m} M_{D}{}^{i\alpha} \ \bar{\nu}_{Ri} \ \nu_{L\alpha} + \frac{1}{2} M_{N}{}^{ij} \ \bar{\nu}_{Ri} \ \nu_{Rj}^{c} + \text{h.c.}$$

$$\nu^{C} \text{ indicates a charge conjugated field } (\nu^{c} = C\bar{\nu}^{T})$$

$$M_{D} \text{ is a complex } m \times 3 \text{ matrix}$$
and M_{N} is a symmetric matrix of dimension $m \times m$

DIRAC NEUTRINOS

 \heartsuit Forcing $M_N=0$ leads to a Dirac mass term which is generated from Yukawa interactions after spontaneous electroweak symmetry breaking

$$Y_{\nu}{}^{i\alpha} \bar{\nu}_{Ri} \bar{\phi}^{\dagger} L_{L\alpha} \Rightarrow M_D{}^{i\alpha} = Y_{\nu}{}^{i\alpha} \frac{v}{\sqrt{2}}$$

similarly to charged fermion masses

- Such a mass term conserves total lepton number but it breaks the lepton flavor number symmetries
- \heartsuit For m=3 we can identify the hypercharge singlets with right-handed component of 4-spinor neutrino field
- \circledcirc Since matrix Y is (in general) a complex 3×3 matrix flavor neutrino fields ν_e,ν_μ and ν_τ do not have a definite mass

MAAJORANA NEUTRINOS

If $M_N \neq 0$ receive important contribution from Majorana mass term

such a term is singlet of SM gauge group

Therefore - it can appear as bare mass term

Furthermore - since it involves two neutrino fields it breaks lepton number by two units

More generally - such a term is allowed only if neutrinos carry no additive conserved charge

This is reason that such terms are not allowed for any charged fermions which (by definition) carry $U(1)_{\rm EM}$

MAJORANA NEUTRINOS (CONT'D)

Dirac mass terms are protected by symmetries of SM (neutrino masses \propto EW symmetry-breaking scale)

Majorana mass term is SM singlet elements of $\,M_N$ are not protected by SM symmetries

It is plausible that Majorana mass term is generated by new physics beyond SM and right-handed chiral neutrino fields ν_R belong to nontrivial multiplets of symmetries of high energy theory





SM accuracy has been shown in a variety of experiments to an amazing level - some observables beyond even one part in a million

Saga of SM is still exhilarating because it leaves all questions of consequence unanswered



Most evident of unanswered questions is why weak interactions are weak





In gauge theory natural values for m_W are zero or Planck mass SM does not contain physics that dictates why its actual value is of order 100 GeV

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THE UGLY



ORDER OF MAGNITUDE OF MASSES OF QUARKS & LEPTONS

THE GOOD

Already in 1934 Fermi provided an answer with a theory that prescribed a quantitative relation between fine structure constant and weak coupling $G_F \sim \alpha/m_W^2$ Although Fermi adjusted m_W to accommodate strength and range of nuclear radioactive decays one can readily obtain a value of m_W of 40 GeV from observed μ decay rate for which proportionality factor is $\pi/\sqrt{2}$ Answer is off by a factor of 2 because discovery of parity violation and neutral currents was in future and introduces an additional factor $1 - m_W^2/m_Z^2$

$$G_F = \left[\frac{\pi\alpha}{\sqrt{2}m_W^2}\right] \left[\frac{1}{1 - m_W^2/m_Z^2}\right] (1 + \Delta r)$$

Fermi could certainly not have anticipated that we now have a renormalizable gauge theory that allows us to calculate radiative correction Δr to his formula

HIGGS MASS RADIATIVE CORRECTIONS

One of most acute problems connected with ultraviolet divergences concerns radiative corrections to mass appearing in Higgs potential $V=\mu^2\Phi\Phi^\dagger+\lambda(\Phi^\dagger\Phi)^2$

f, W, and H self-coupling radiative corrections to Higgs mass are

VELTRAAN CONDITION

Difference between bare and renormalized masses is

$$\Delta \mu^2 = \frac{1}{64\pi^2} \left(9g^2 + 3g'^2 + 24\lambda - 8\sum_f N_f Y_f^2 \right) \Lambda^2$$
$$\simeq \frac{3}{16\pi^2 v^2} \left(2m_W^2 + m_Z^2 + m_H^2 - 4m_t^2 \right) \Lambda^2$$

SM works amazingly well by fixing Λ at electroweak scale generally assumed \blacktriangleright this indicates existence of new physics beyond SM Following Weinberg

$$\mathcal{L}(m_{\rm W}) = |\mu^2| H^{\dagger} H + \frac{1}{4} \lambda (H^{\dagger} H)^2 + \mathcal{L}_{\rm SM}^{\rm gauge} + \mathcal{L}_{\rm SM}^{\rm Yukawa} + \frac{1}{\Lambda} \mathcal{L}^5 + \frac{1}{\Lambda^2} \mathcal{L}^6 + \dots$$

where operators of higher dimension parametrize physics beyond SM Some have resorted to rather extreme lengths by proposing that factor multiplying unruly quadratic correction $(2m_W^2 + m_Z^2 + m_H^2 - 4m_t^2)$ must vanish EXACTLY!

By most conservative estimates this new physics is within our reach

NEW PHYSICS AT THE TEV SCALE?

Avoiding fine tuning requires $\Lambda \lesssim 2-3$ to be revealed by CERN LHC E.G. for $m_H = 115-200~{\rm GeV},$

$$\left|\frac{\Delta\mu^2}{\mu^2}\right| = \frac{\delta v^2}{v^2} \le 10 \Rightarrow \Lambda = 2 - 3 \text{ TeV}$$

we have implicity used $v^2=-\mu^2/\lambda$ [valid in approximation disregarding terms beyond ${\cal O}(H^4)$ in Higgs potential] Let's contemplate possibilities

Veltman condition happens to be satisfied and this would leave particle physics with an ugly fine tuning problem

 This is very unlikely - LHC must reveal Higgs physics already observed via radiative correction or at least discover physics that implements Veltman condition
 It must appear at 2 - 3TeV even though higher scales can be rationalized when accommodating selected experiments

SUSY

Supersymmetry predicts that interactions exist that would change fermions into bosons and vice versa and that all known fermions (bosons) have SUSY boson (fermion) partner

SUSY offers solution of bad behavior of radiative corrections in SM

As for every boson there is a companion fermion bad divergence associated with Higgs loop cancelled by a fermion loop with opposite sign



even though it elegantly controls quadratic divergence by cancellation of boson and fermion contributions it is already fine-tuned at a scale of $2-3{
m TeV}$

THE BAD ...



Happy Chanksgiving

